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# Isometric Embeddings in Binary Black Hole Merger Simulations

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# Isometric Embeddings

# The Isometric Embedding Problem

- Consists in visualizing a surface in 3D Euclidean space that would have the same metric as a given geometry.
- Embedding equations:

$$ds^2 = g_{\theta\theta}\, d\theta^2 + g_{\phi\phi}\, d\phi^2 + 2g_{\theta\phi}\, d\theta d\phi \tag{1}$$

$$ds^2 = dx^2 + dy^2 + dz^2 \tag{2}$$

$$\Rightarrow \begin{cases} (\partial_\theta x)^2 + (\partial_\theta y)^2 + (\partial_\theta z)^2 = g_{\theta\theta} \\ (\partial_\theta x)(\partial_\phi x) + (\partial_\theta y)(\partial_\phi y) + (\partial_\theta z)(\partial_\phi z) = g_{\theta\phi} \\ (\partial_\phi x)^2 + (\partial_\phi y)^2 + (\partial_\phi z)^2 = g_{\phi\phi} \end{cases} \tag{3}$$

- Known to be difficult to solve in general since the foundations of differential geometry. Being quadratic in the highest-order derivatives, equation (3) cannot be solved with usual PDE algorithms.

# Motivations

- Remove arbitrariness when visualizing apparent horizons.
  - Current visualizations depend on the choice of coordinates.
- Useful for computing quasilocal quantities such as mass, energy, and angular momentum.

# Remarks

- Not all 2-geometries are embeddable.
  - Famous example: Kerr black hole with dimensionless spin $\chi = a/M > \sqrt{3}/2$.
- Embeddings are unique only up to translational and rotational degrees of freedom in Euclidean space.

# Numerical Methods

# Matrix Method

- Adapted from Tichy et al.[1]
- Pseudospectral collocation: represent the embedding functions $x^i \in \{x, y, z\}$ as an expansion of basis functions.

$$x^i(\theta, \phi) = d^{in} D_n(\theta, \phi) \tag{4}$$

$$\Rightarrow \partial_A x^i(\theta, \phi) = d^{in} \partial_A D_n(\theta, \phi) \tag{5}$$

  $D_n(\theta, \phi) \longleftrightarrow Y_{\ell,m}(\theta, \phi)$

- Solve for the coefficients $d^{in}$ using Newton-Raphson iterations. At each step, we find $\delta d^{in}$ using LAPACK's `dgelsy` routine.[2]

[1] Wolfgang Tichy *et al* 2015 Class. Quantum Grav. **32** 015002

[2] This least-squares matrix solver scales poorly with resolution.

# Relaxation Method

- Idea: find the dyad vectors $\vec{e}_A \in \{\vec{e}_\theta, \vec{e}_\phi\}$, which are the basis vectors of $A \in \{\theta, \phi\}$.

$$\partial_A \vec{x} = \vec{e}_A \tag{6}$$

- To be a valid coordinate basis, $\vec{e}_\theta$ and $\vec{e}_\phi$ need to commute.

$$[\vec{e}_\theta, \vec{e}_\phi] := \partial_\theta \vec{e}_\phi - \partial_\phi \vec{e}_\theta \to 0 \tag{7}$$

- With (6), the embedding equations become

$$\vec{e}_A \cdot \vec{e}_B = g_{AB}. \tag{8}$$

- Approach: rotate $\vec{e}_A$ to achieve (7) and damp any violations of the constraints (8).
- Once we have $\vec{e}_A$, we solve (6) for $\vec{x}$.

# Relaxation Method (continued)

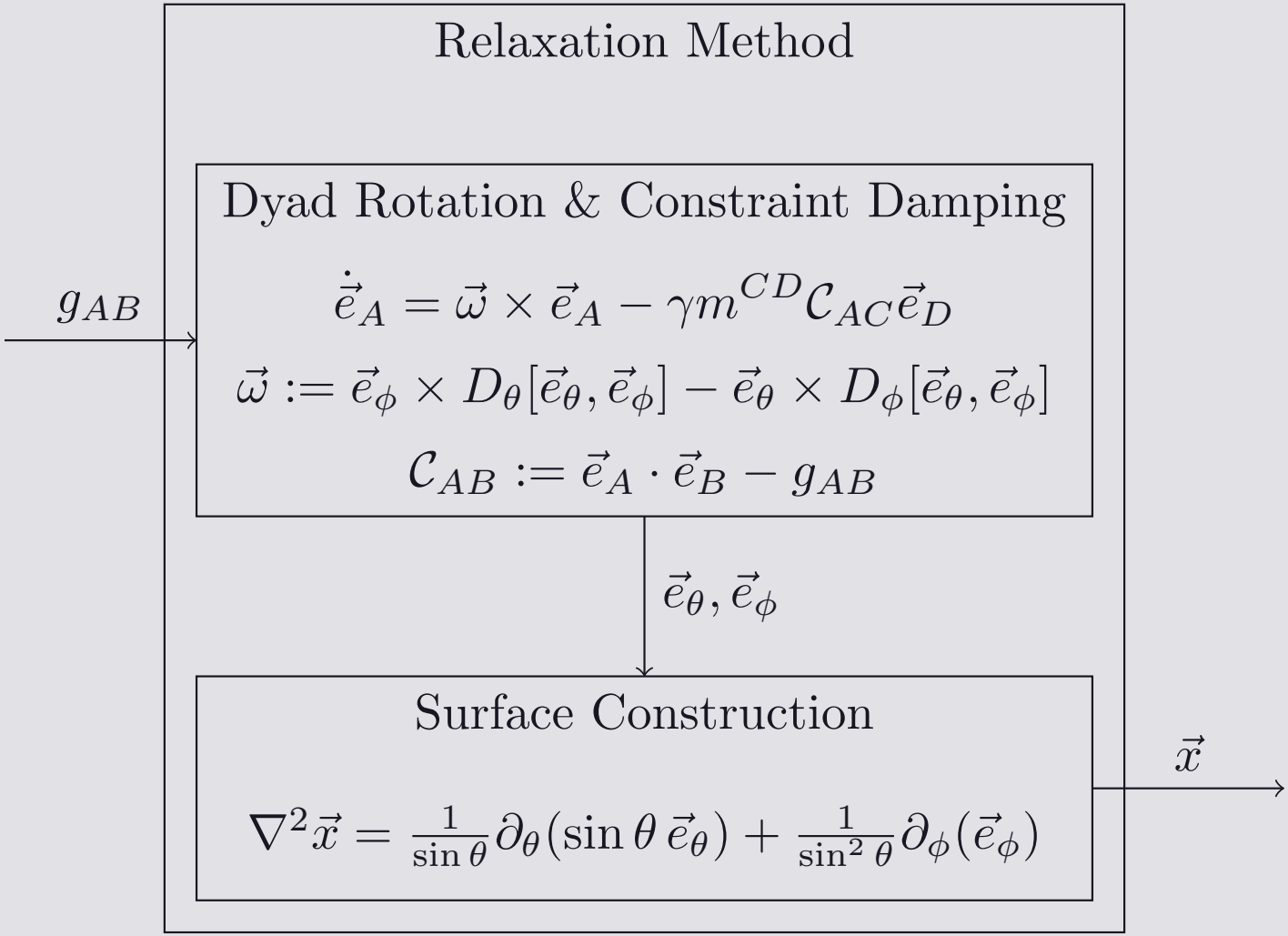



**Figure 1:** Summary of the relaxation method.

# Code Implementations

- Both methods were implemented on the Spectral Einstein Code (`SpEC`), which was used for all results presented here.
- As the relaxation method is a novel algorithm, we initially wanted to test it separately. Then, we created the Finite Difference Embedding Code (`FiDEC`), which is open-source.[3]

[3] Link to `FiDEC` is at the end of the presentation.

# Test Cases

# Selected Test Surfaces

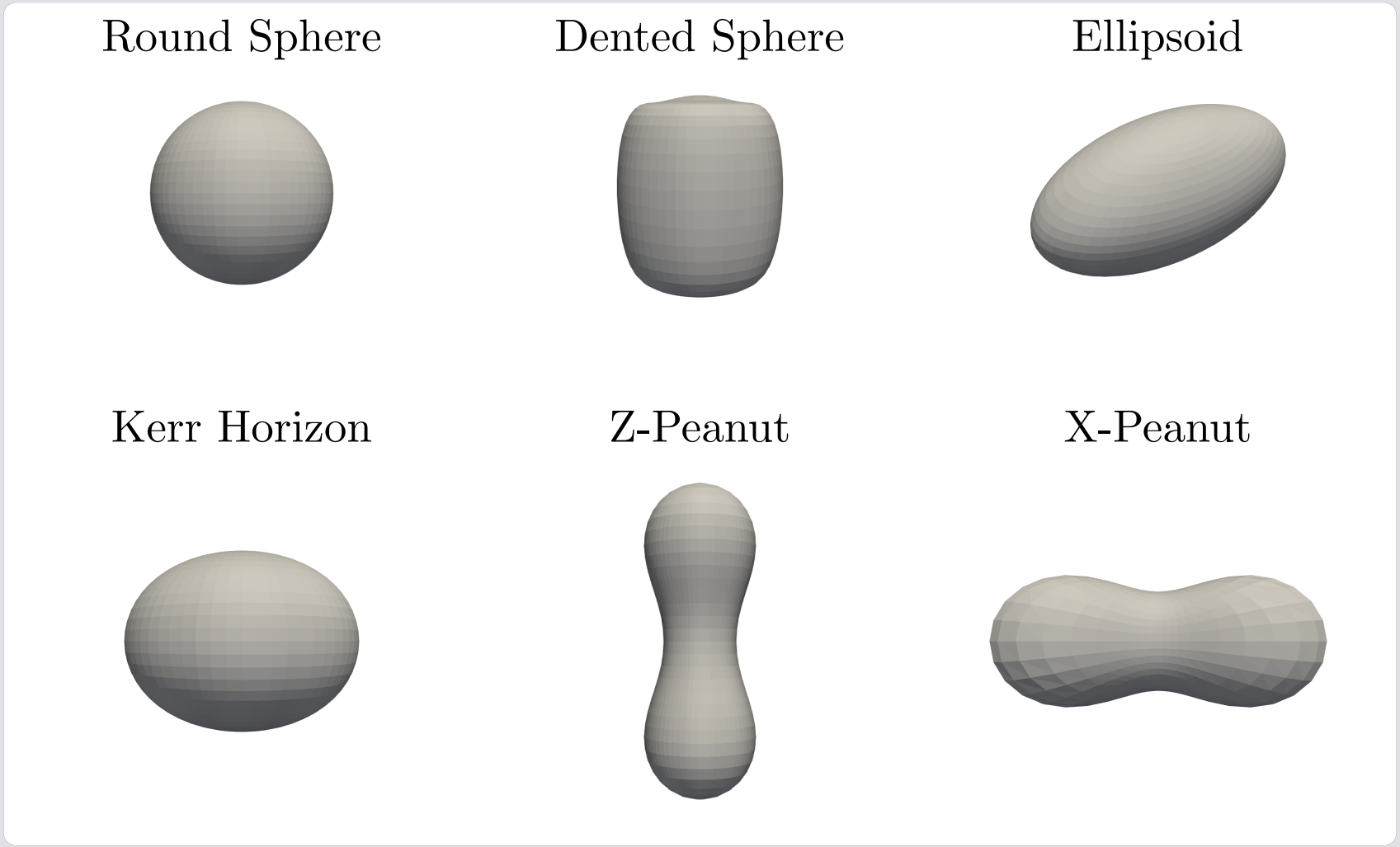



**Figure 2:** Selected surfaces used as test cases for the relaxation and matrix methods.

# Convergence

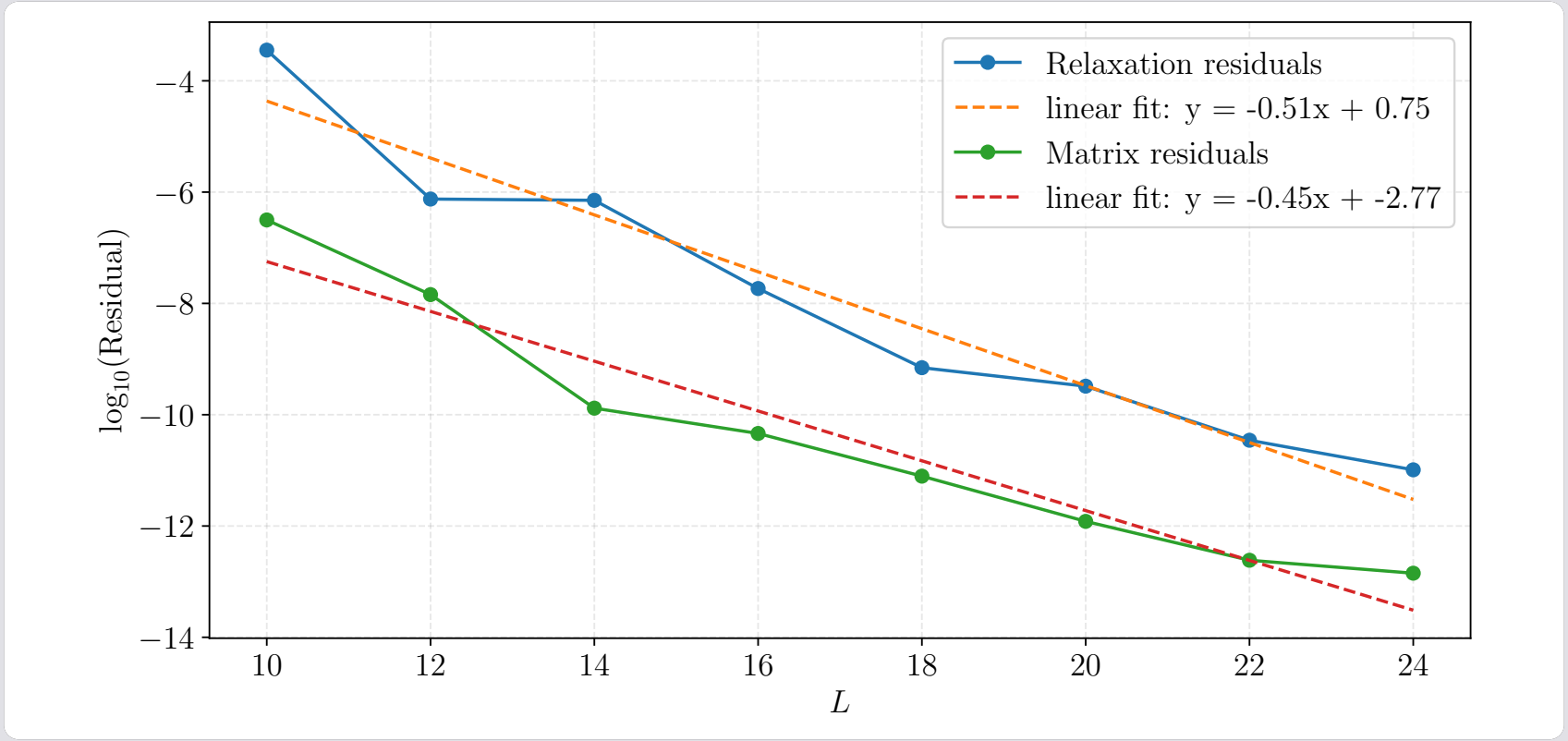



**Figure 3:** Convergence plot[4] for a Kerr horizon with $\chi = 0.7$.

[4]We define our residuals as $R_{AB} := (\partial_A \vec{x}) \cdot (\partial_B \vec{x}) - g_{AB}$, from which we get a norm and average over the surface to get a scalar.

# Time Scaling

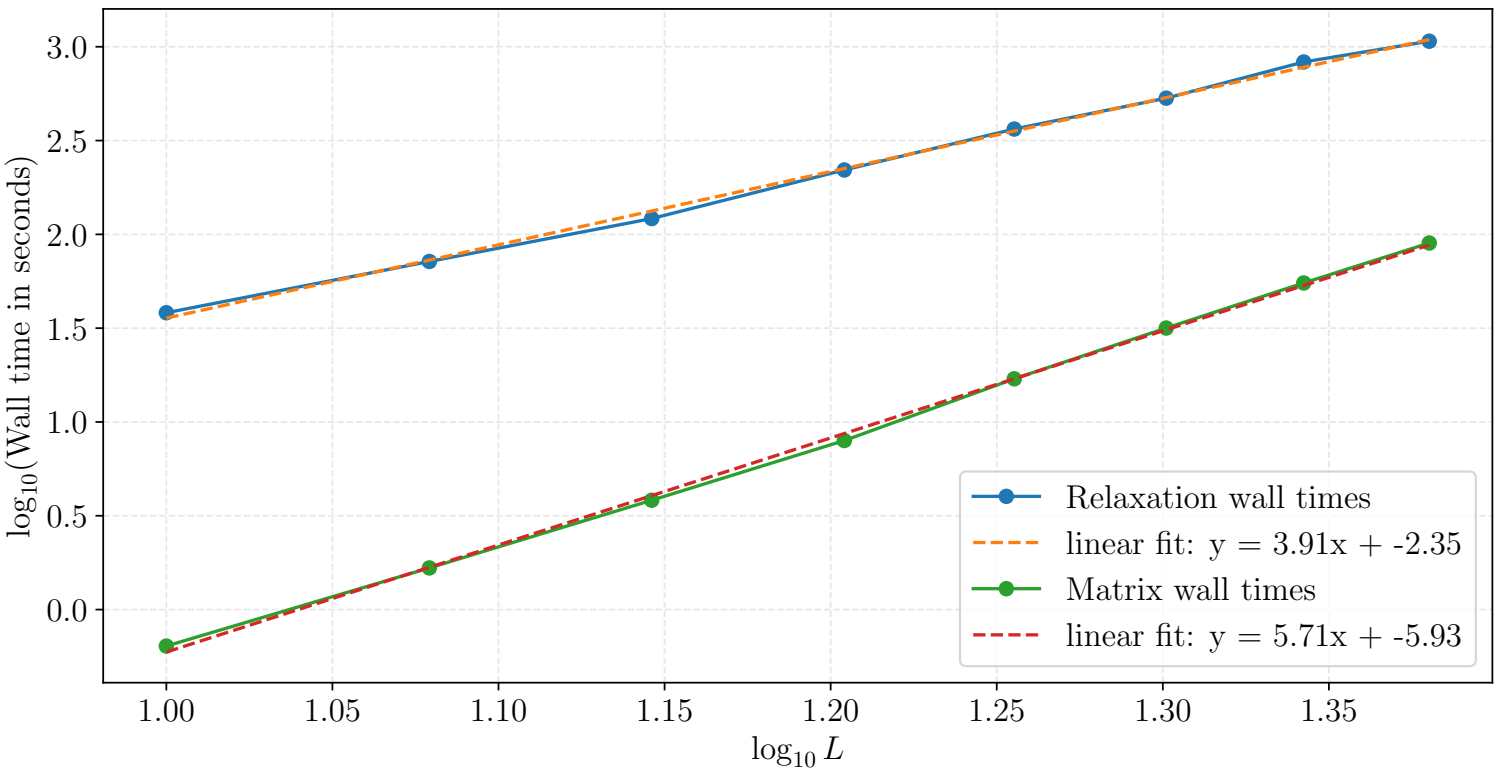



**Figure 4:** Time scaling plot for a Kerr horizon with $\chi = 0.7$. The relaxation method scales as $\sim O(L^4)$, whereas the matrix method scales as $\sim O(L^6)$.

# Kerr Embeddability Limit

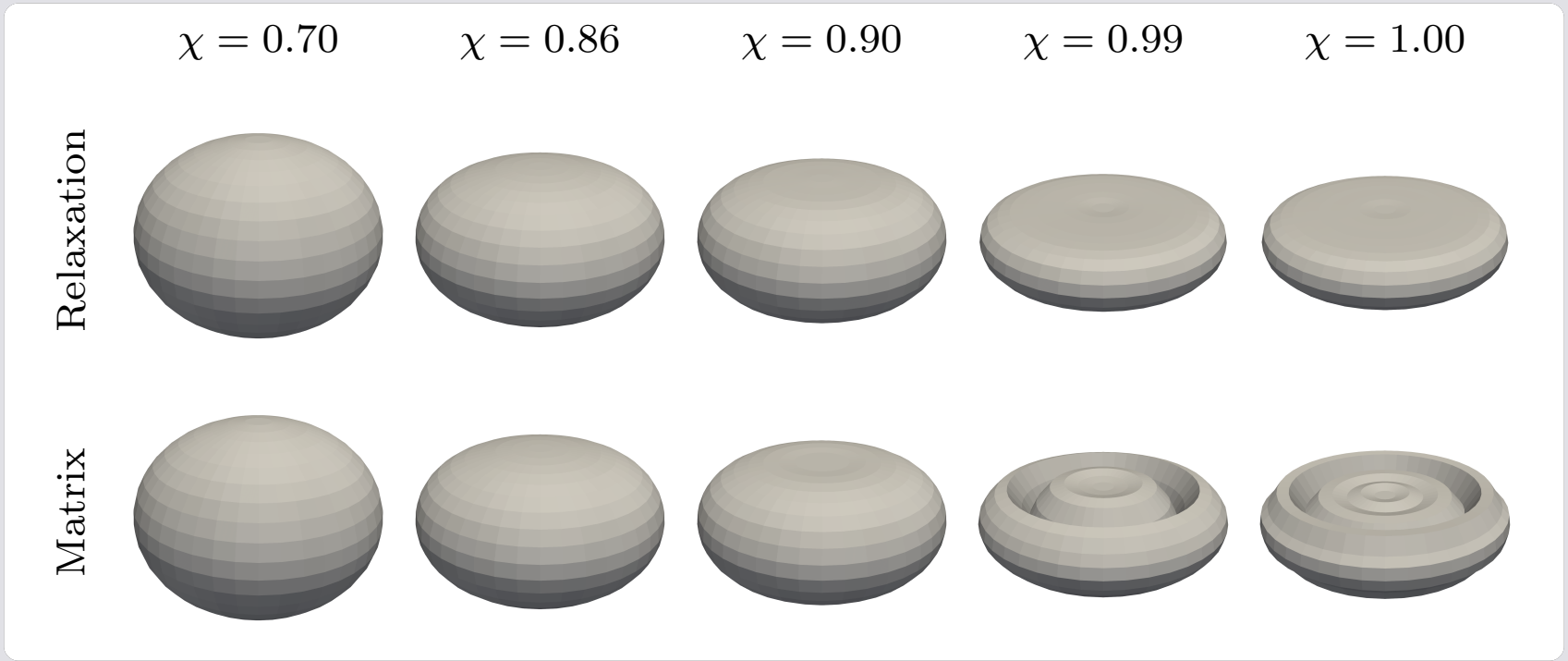



**Figure 5:** Comparison between embedding results from the relaxation and matrix methods as we approach and exceed the embeddability limit $\sqrt{3}/2 \approx 0.866$.

# Kerr Embeddability Limit (continued)

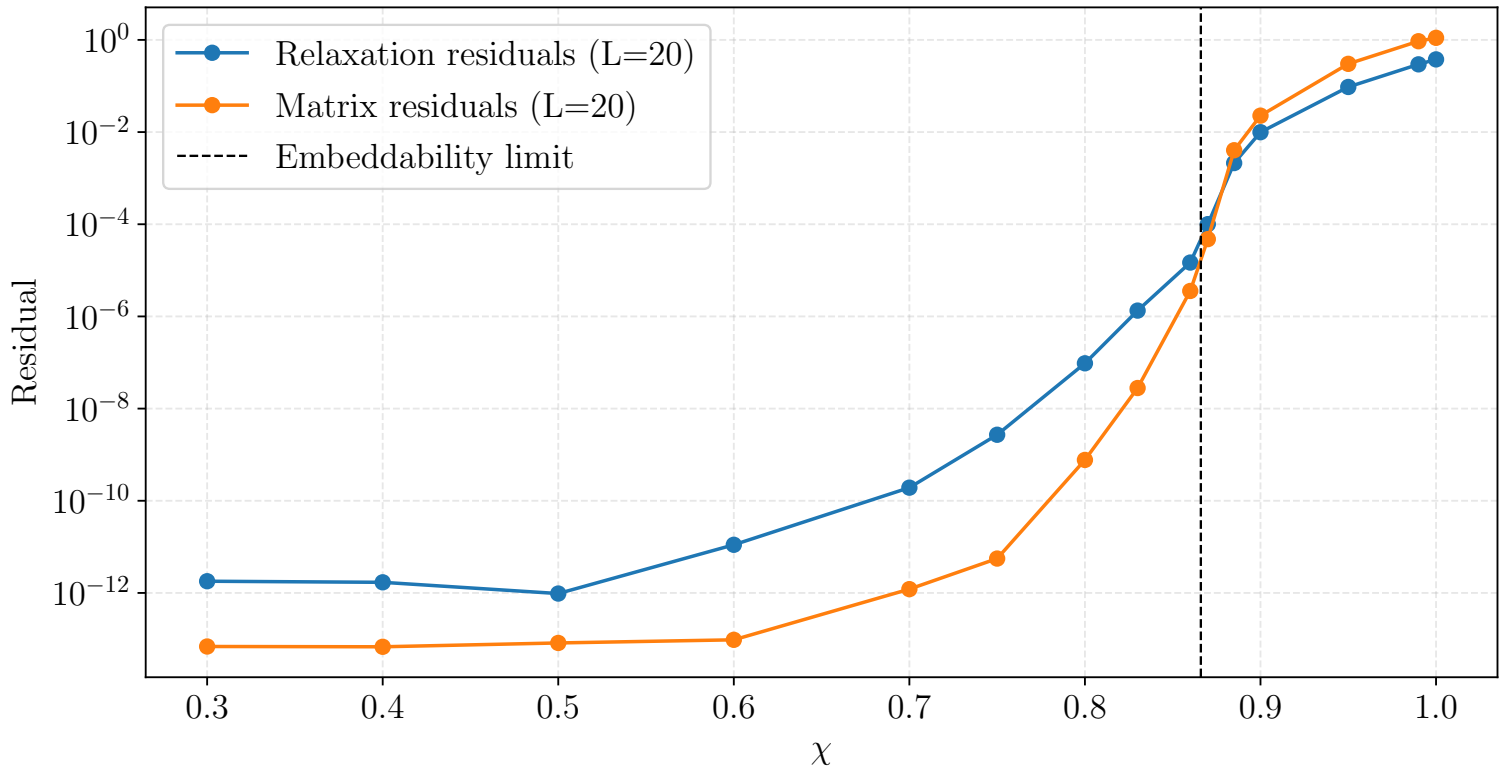



**Figure 6:** Residuals of the relaxation and matrix methods as we vary the dimensionless spin $\chi$ past the embeddability limit $\sqrt{3}/2 \approx 0.866$ for a resolution of $L = 20$.

# Kerr Embeddability Limit (continued)

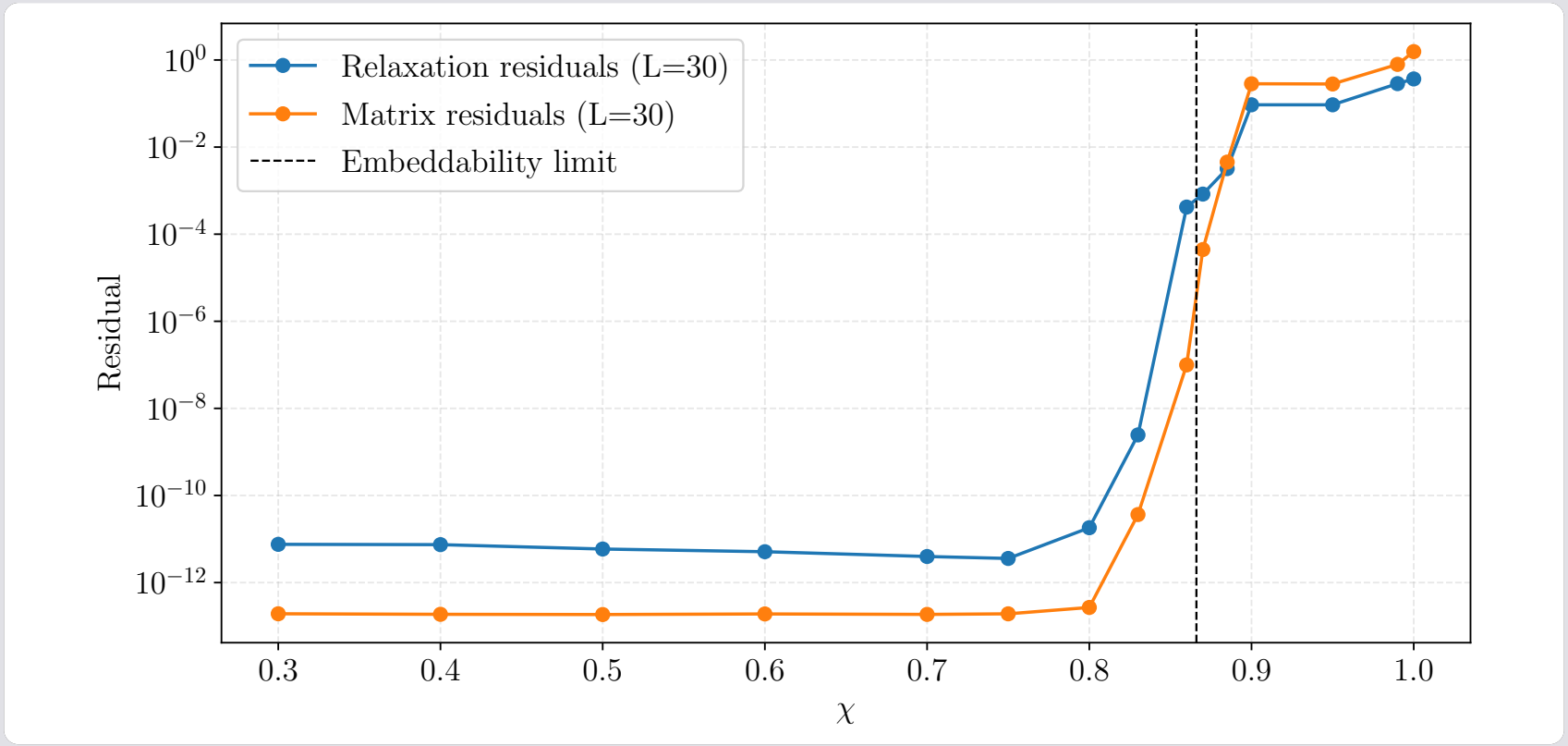
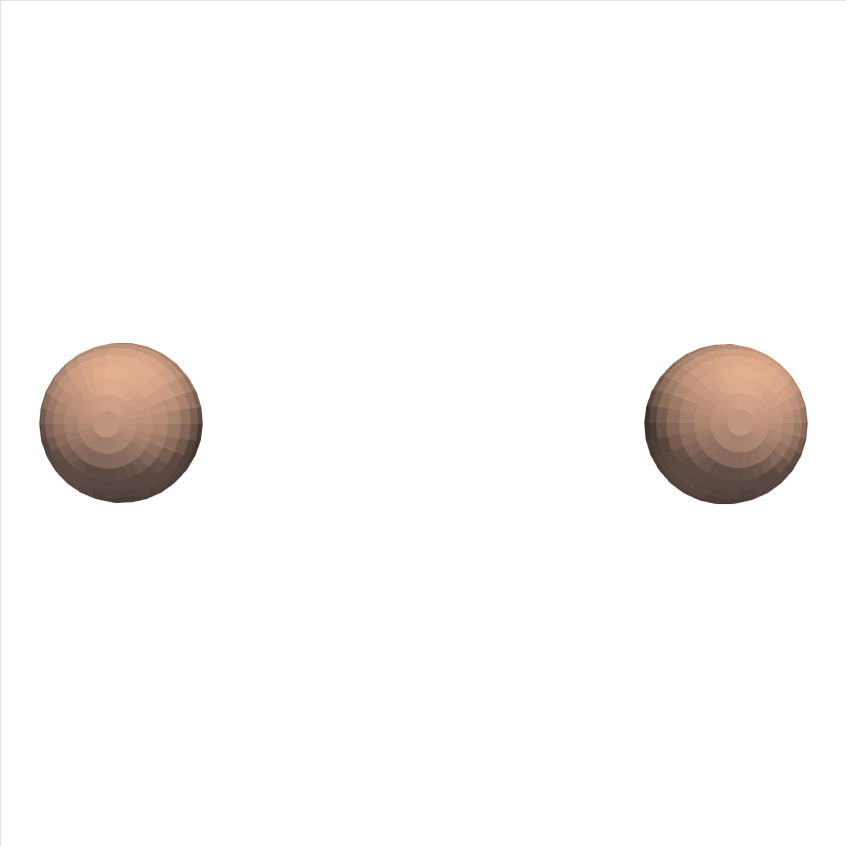



**Figure 7:** Residuals of the relaxation and matrix methods as we vary the dimensionless spin $\chi$ past the embeddability limit $\sqrt{3}/2 \approx 0.866$ for a resolution of $L = 30$.

# BBH Merger Simulations

# Equal-mass Non-spinning $q = 1;\ \chi_A = \chi_B = (0, 0, 0)$



**Video 1:** Evolution of isometric embeddings.

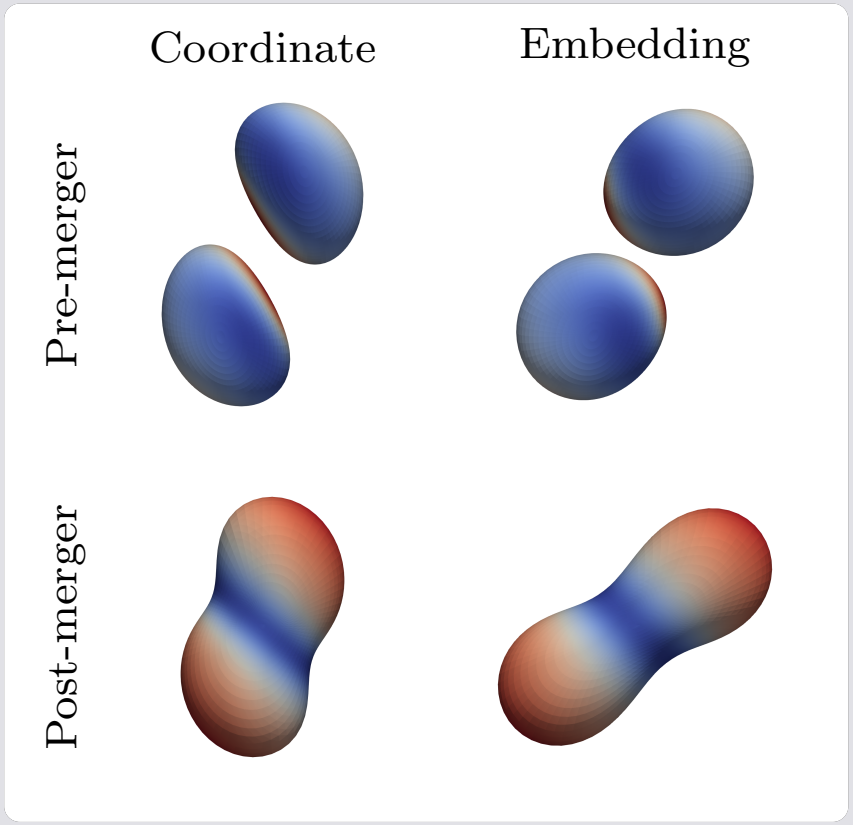
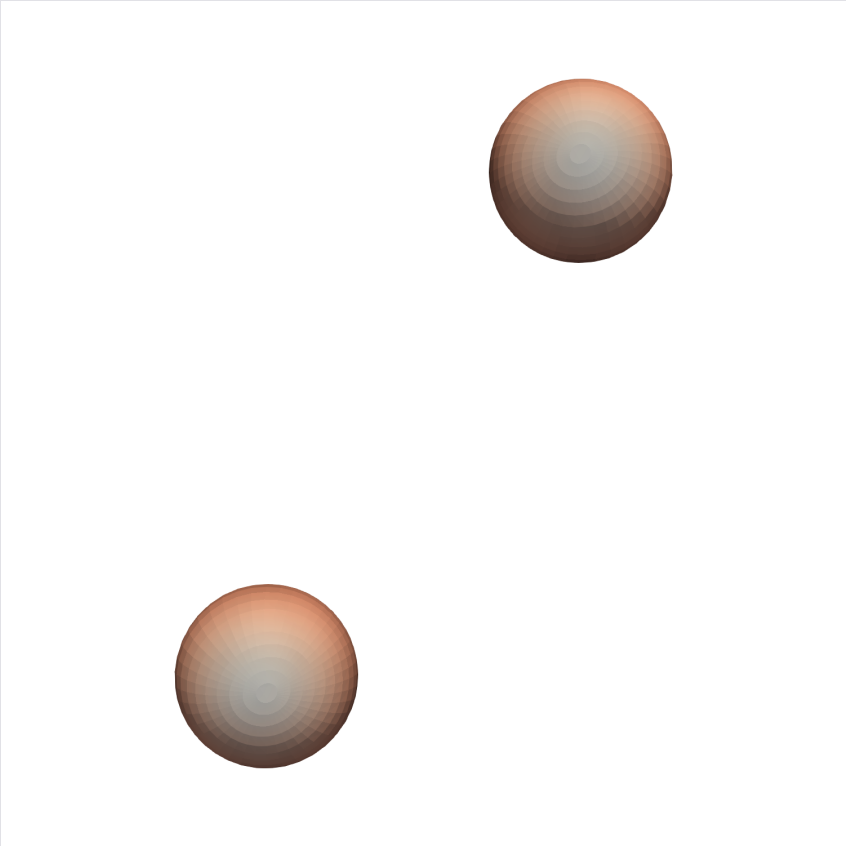



**Figure 8:** Coordinate shapes vs. isometric embeddings.

# Co-aligned

$q = 1;\ \chi_A = \chi_B = (0, 0, 0.6)$



**Video 2:** Evolution of isometric embeddings.

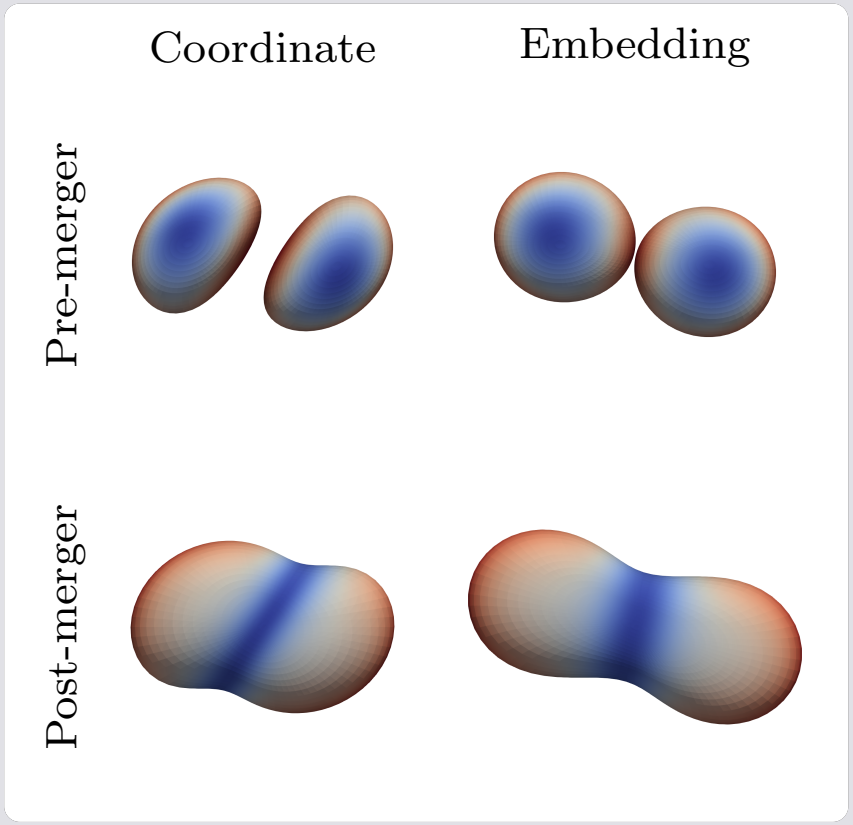
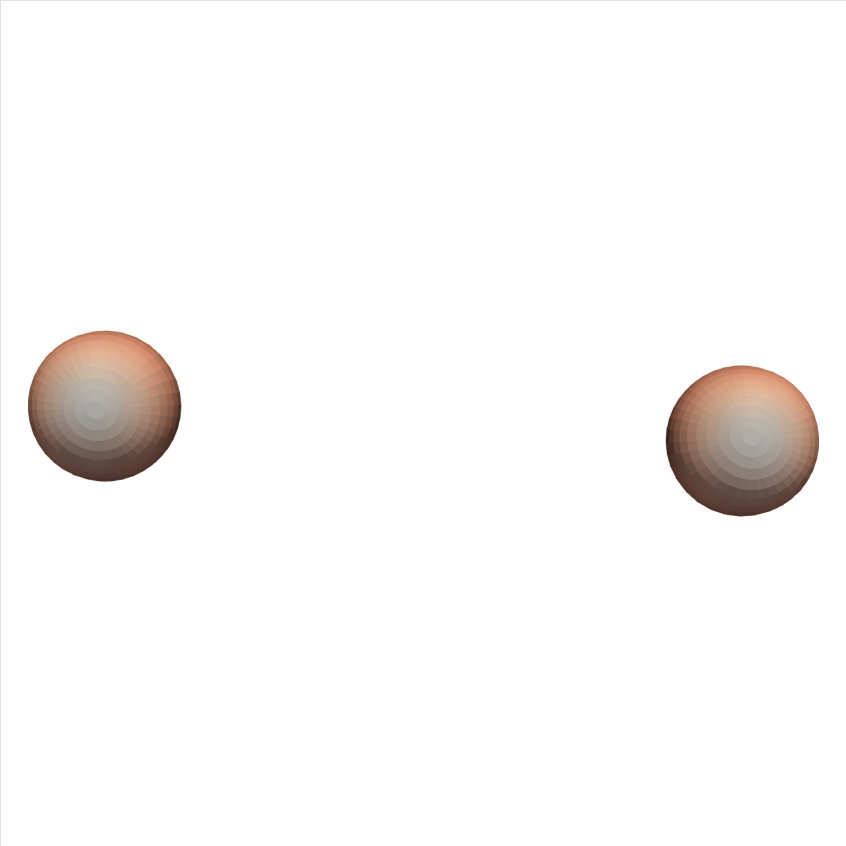



**Figure 9:** Coordinate shapes vs. isometric embeddings.

# Anti-aligned

$q = 4;\ \chi_A = \chi_B = (0, 0, -0.6)$



**Video 3:** Evolution of isometric embeddings.

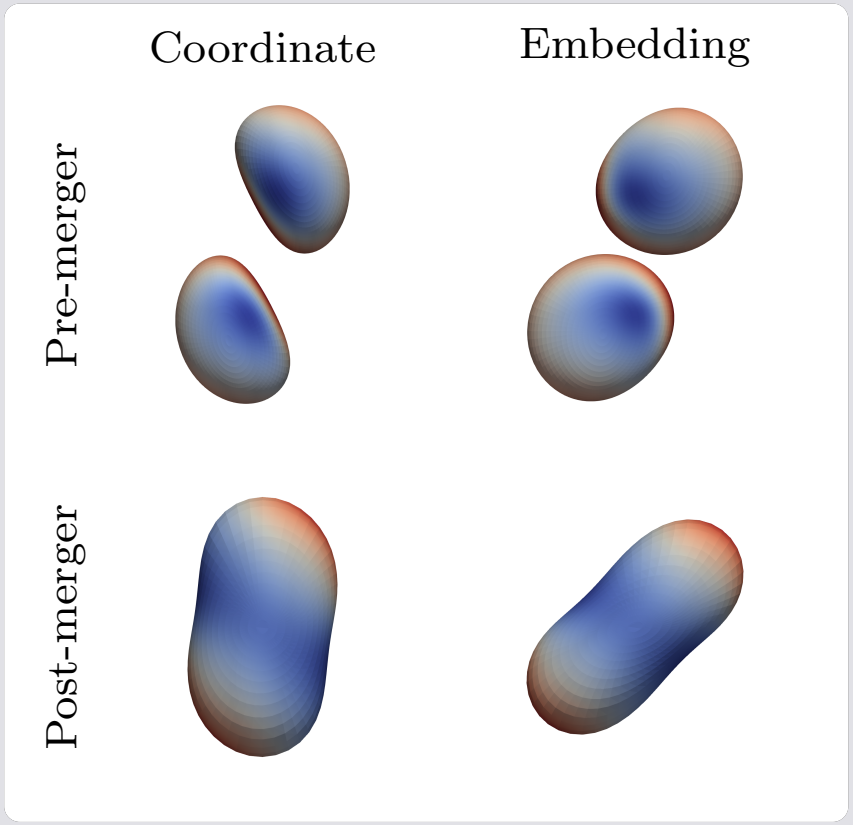
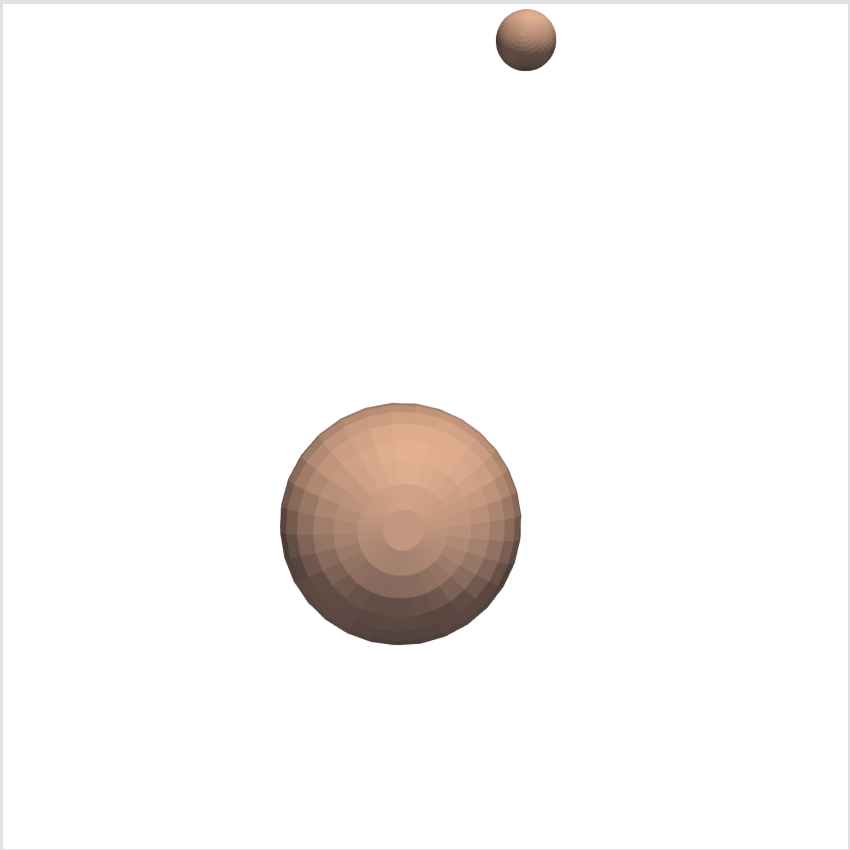



**Figure 10:** Coordinate shapes vs. isometric embeddings.

# Nontrivial Mass Ratio

$q = 4;\ \chi_A = \chi_B = (0, 0, 0)$



**Video 4:** Evolution of isometric embeddings.

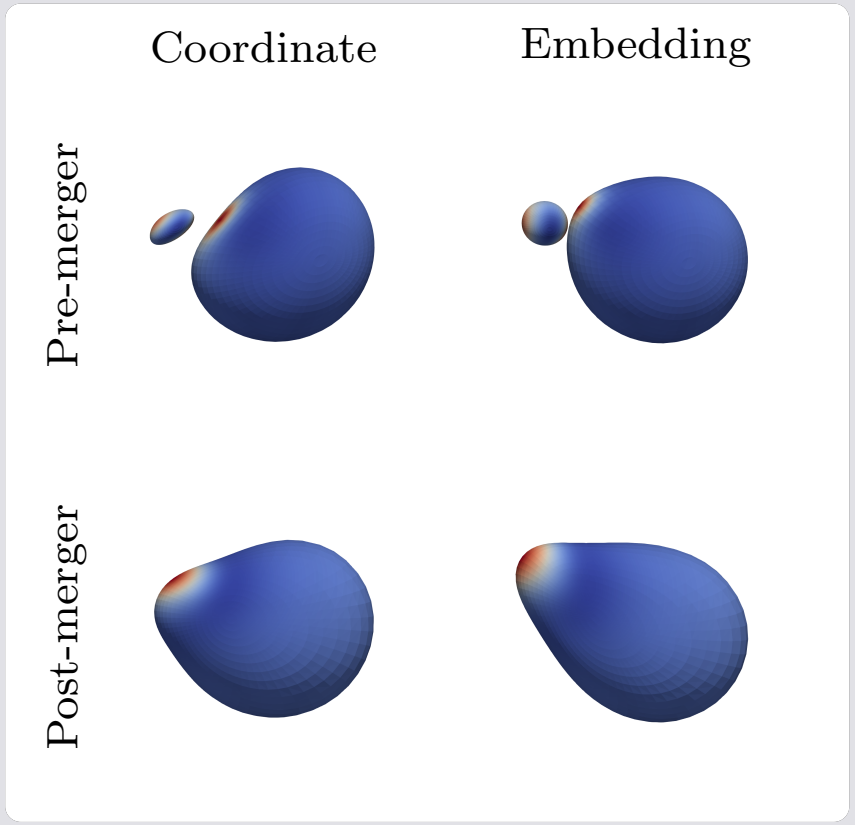
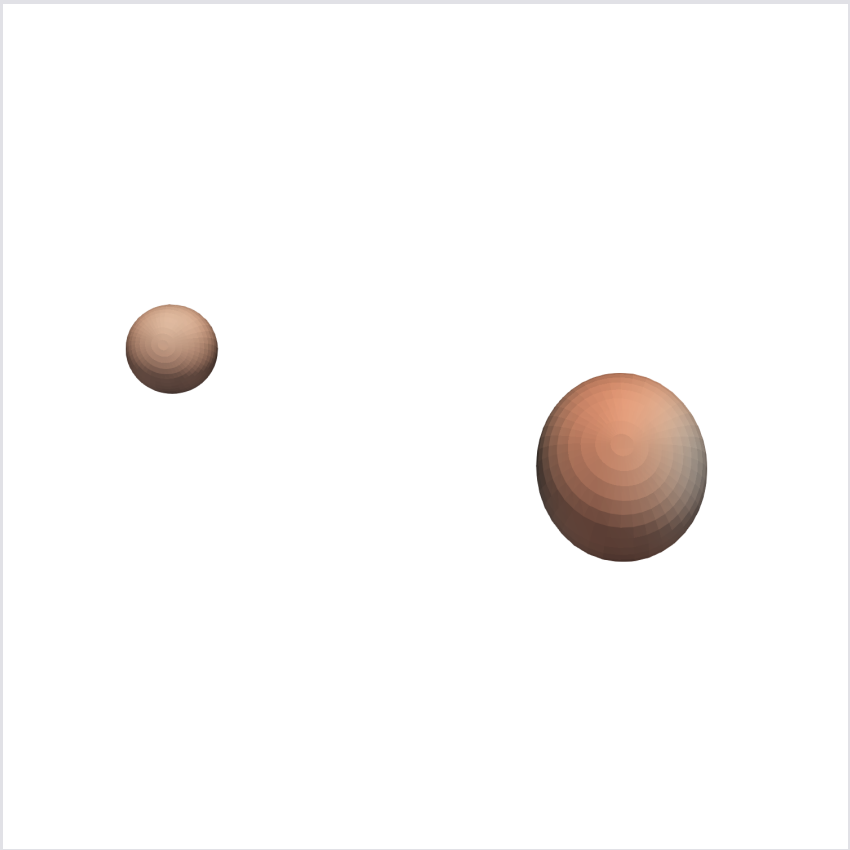



**Figure 11:** Coordinate shapes vs. isometric embeddings.



**Video 5:** Evolution of isometric embeddings.

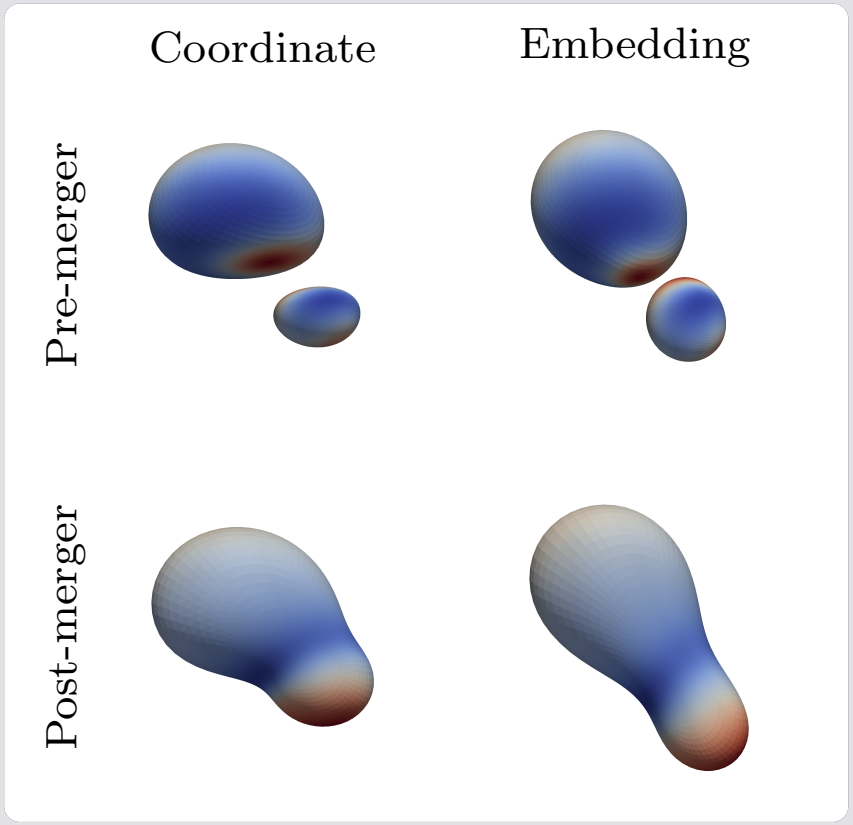



**Figure 12:** Coordinate shapes vs. isometric embeddings.

# Non-embeddability?

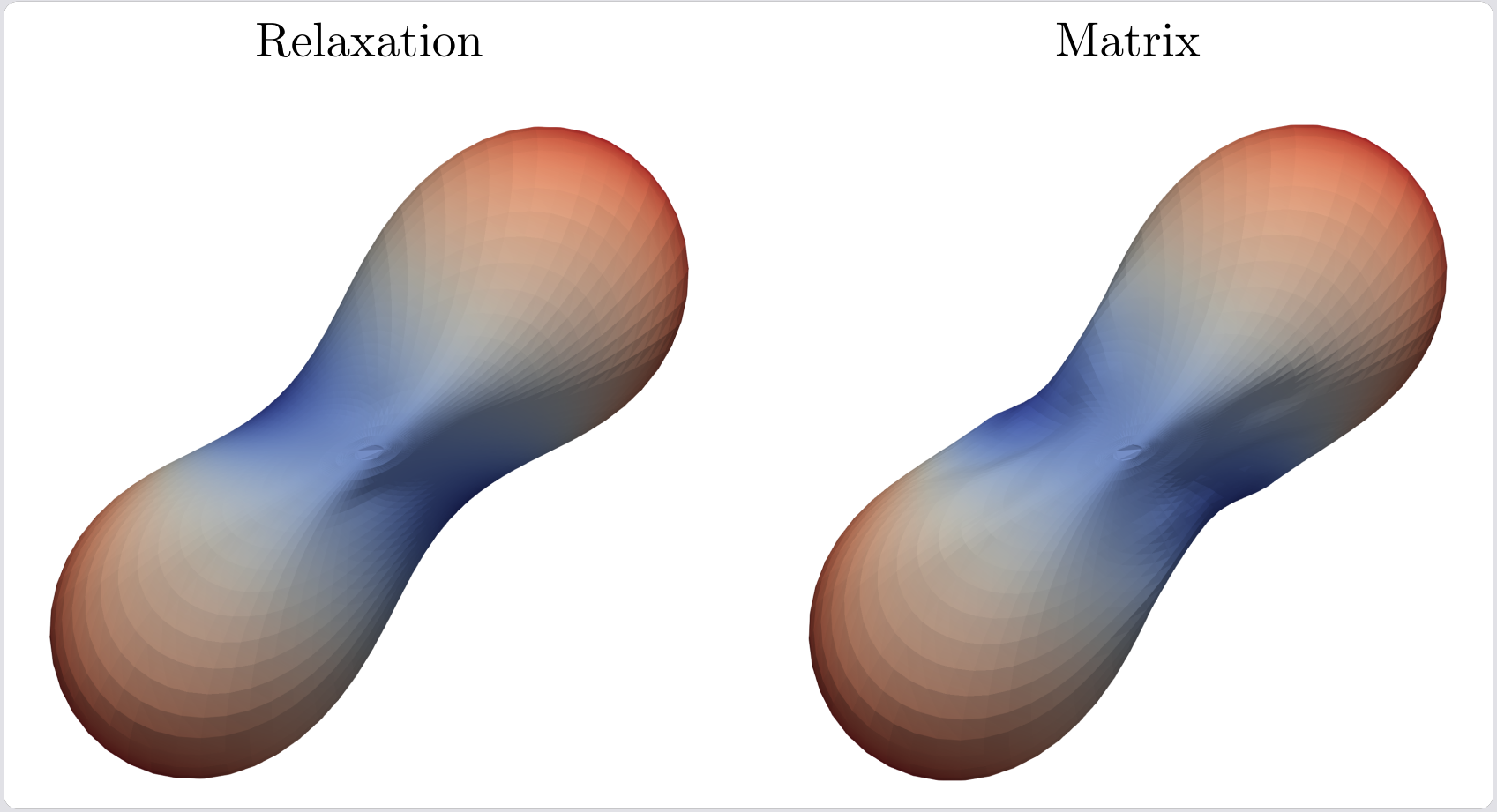



**Figure 13:** Failed embedding results from the relaxation method versus the matrix method for the anti-aligned run.

# Non-embeddability? (continued)

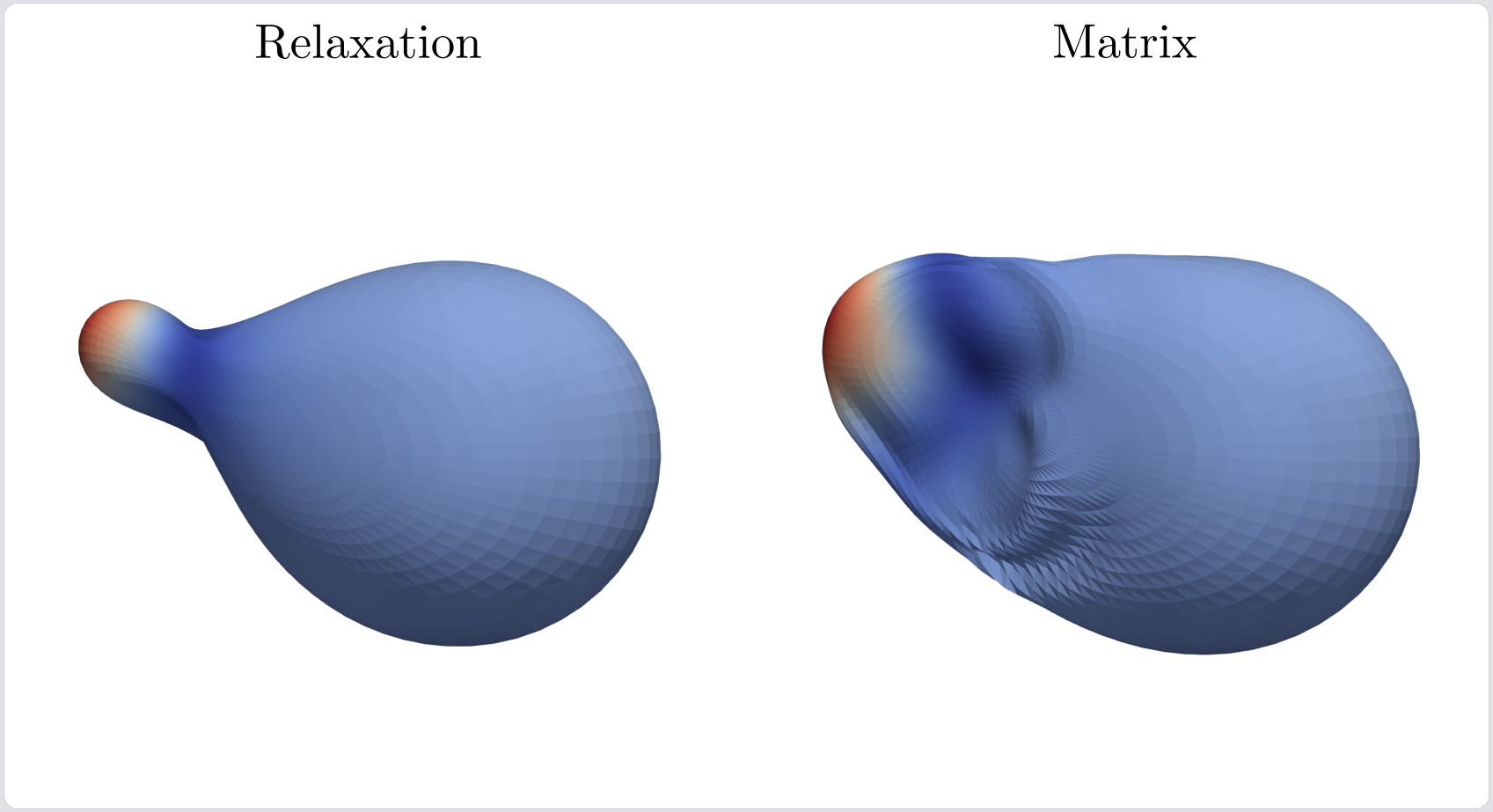



**Figure 14:** Failed embedding results from the relaxation method versus the matrix method for the $q = 4$ run.

# Conclusion

# Summary

- We have developed two efficient numerical methods for the isometric embedding problem in `SpEC`.
- Through test cases, we confirmed that they converge and analyzed their behavior.
- We applied these methods to numerical horizons in BBH merger simulations.

# Future Work

- **Critical embedding**: understand how numerical horizons become embeddable or non-embeddable.
- **Wang-Yau quasilocal mass**: generalize our methods to Minkowski space.[5]
- **Gauge conditions**: use embedding results to choose gauge conditions that lead to more round apparent horizons (and faster mergers).

[5] Mu-Tao Wang and Shing-Tung Yau. "Quasilocal mass in general relativity". *Phys. Rev. Lett.*, 102:021101, Jan 2009.

# Thank you! Any Questions?

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**More Details**

iagomendes.com/embedding

# Extra Slides

# Wall Times

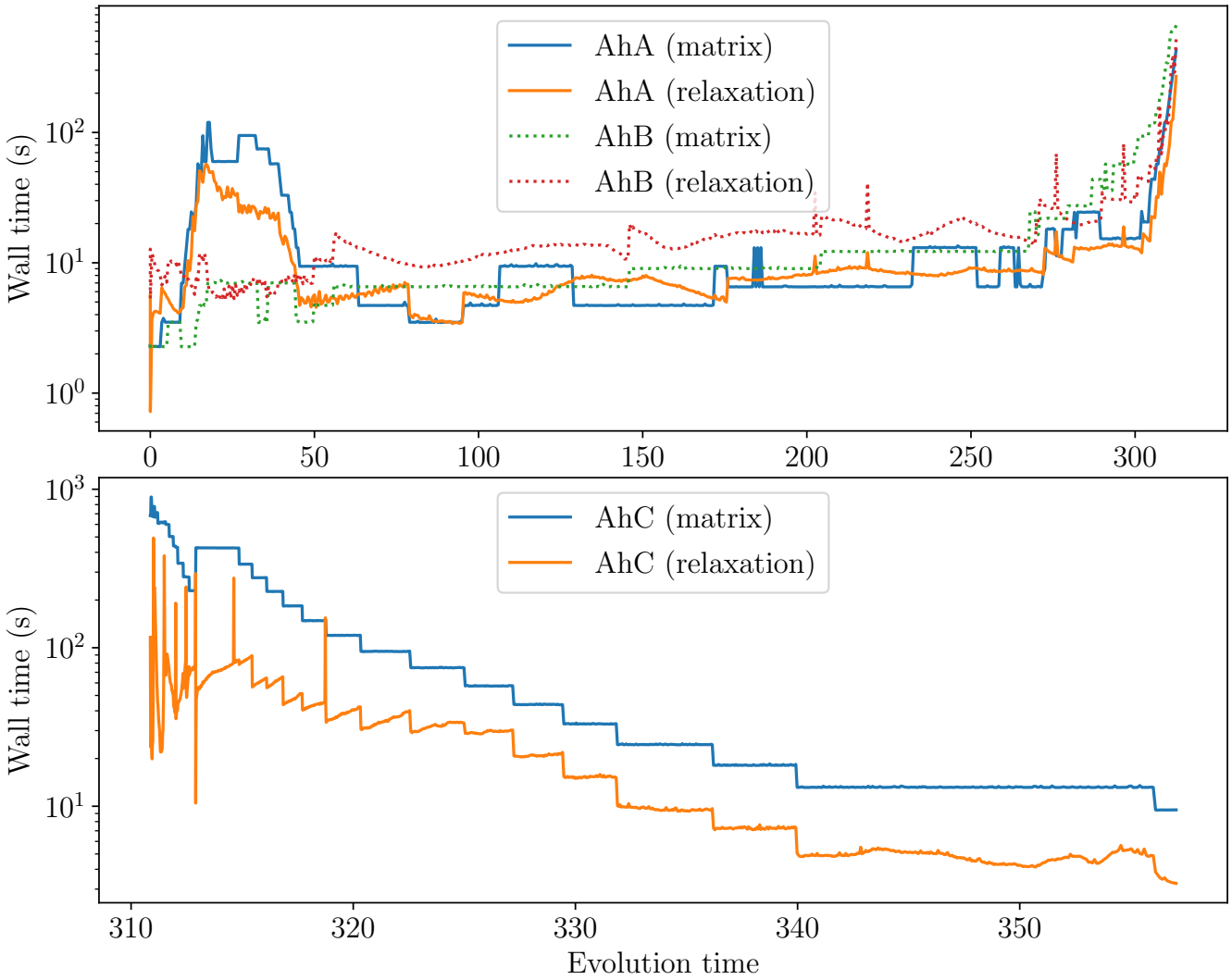


**Figure 15:** Generic run.

# Wall Times (continued)

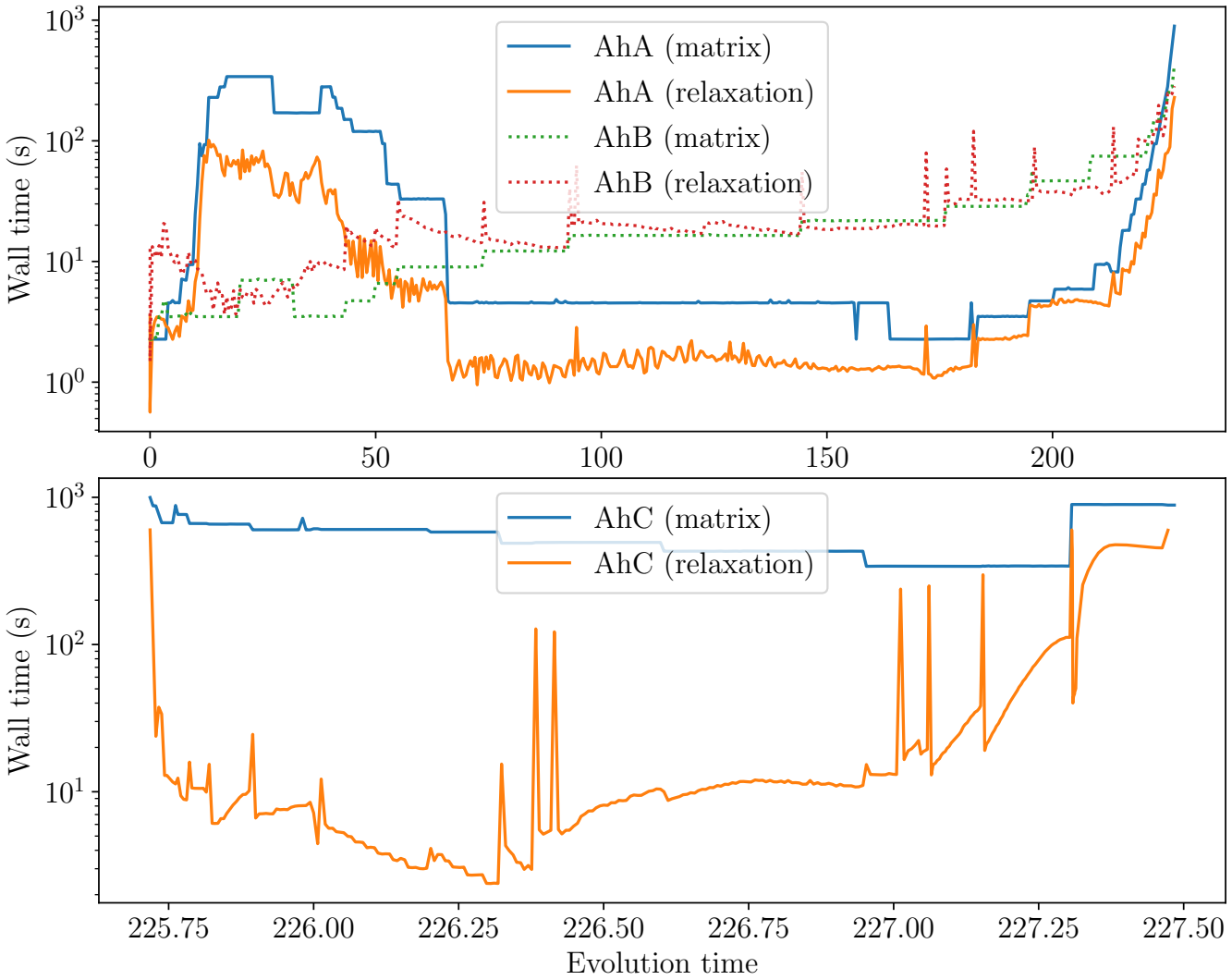


**Figure 16:** $q = 4$ run.

# Residual Evolution

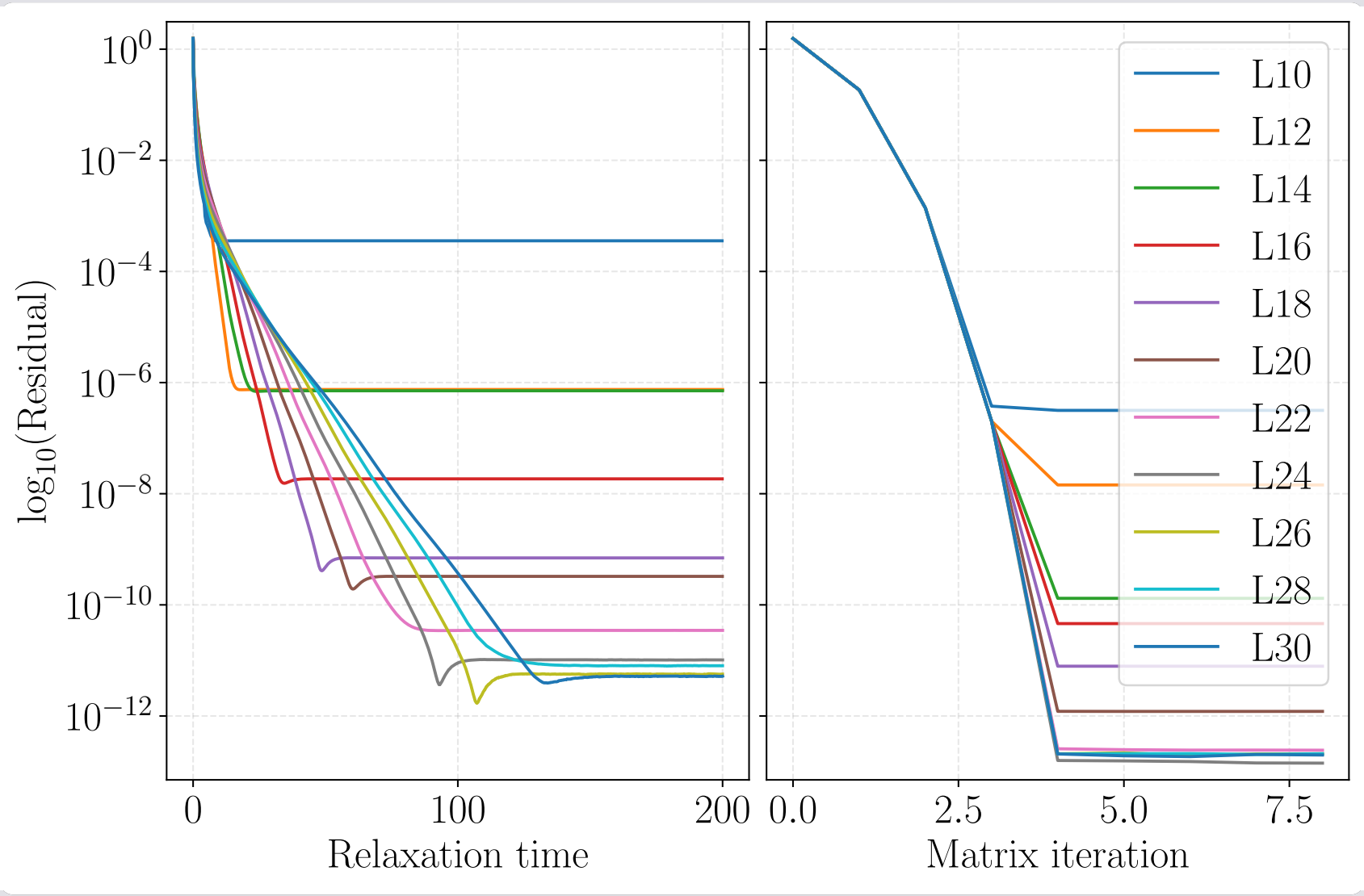


**Figure 17:** Kerr Horizon.

# Relaxation Residual Distribution

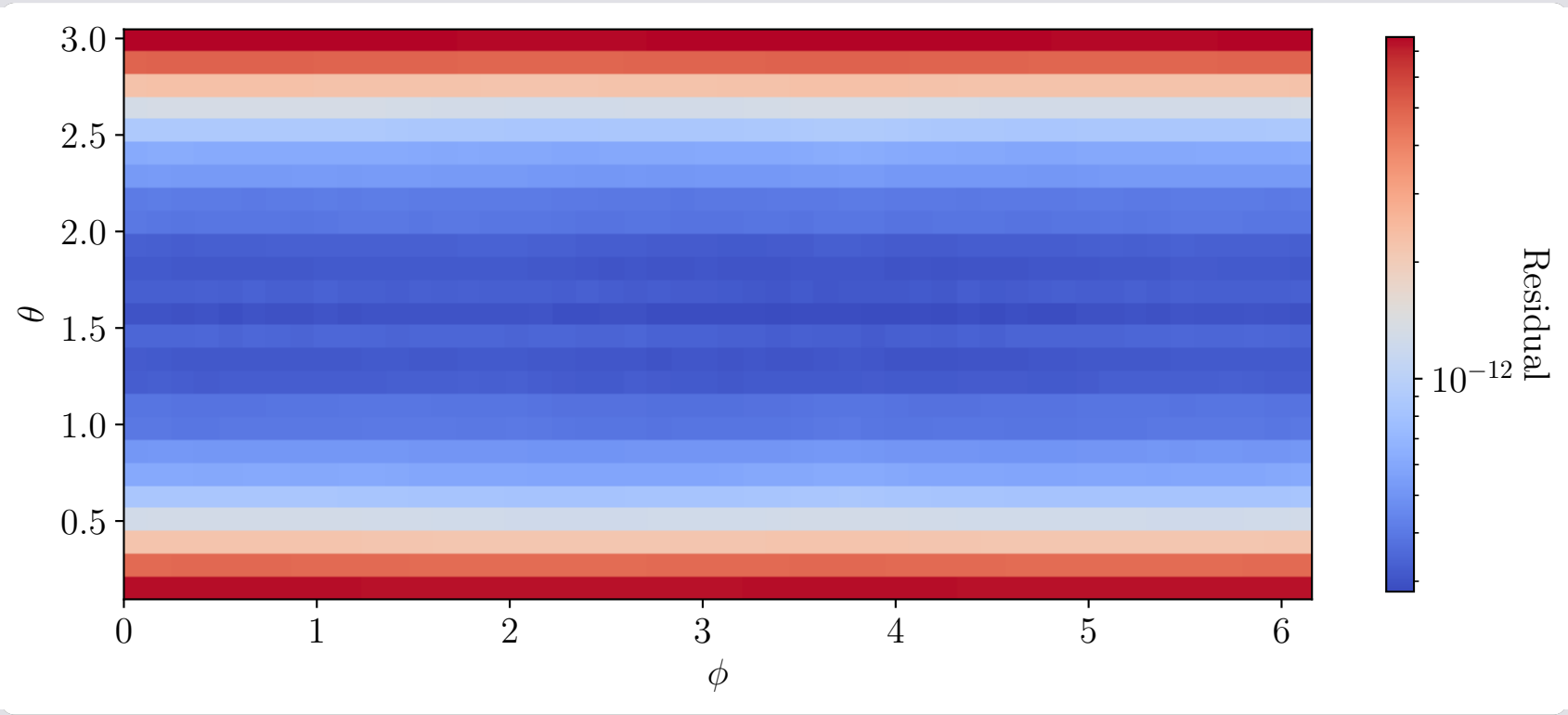


**Figure 18:** Kerr Horizon.